

Mixture of noises and sampling non-log-concave posterior distributions – EUSIPCO 2022 –

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Real life inverse problems

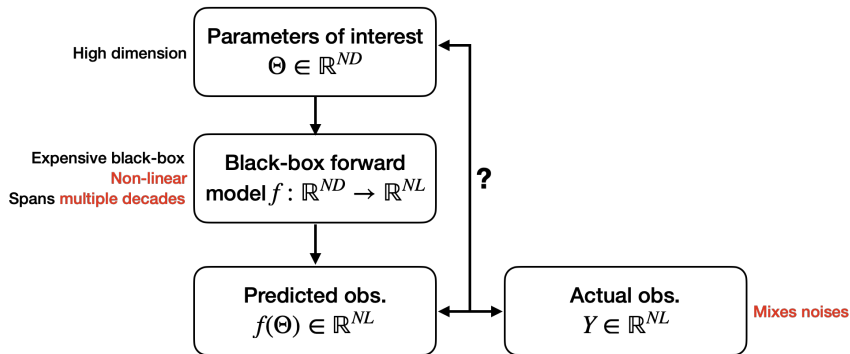


Considered general inverse problem

	Log-likelihood	Proposition
Mixture of noises	Untractable	Approx.
Forward model:		
→ expensive black-box	-	Model reduction
→ Non-linear	Non-concave	MTM kernel
→ Spans multiple decades	Non-grad-Lipschitz	P-MALA kernel
Uncertainty quantification	-	MCMC

Limitation: Sampler restricted to smooth log-posterior because of P-MALA (see slide 9)

Considered general inverse problem



Observation model

$$y_{n,\ell} = \max \left\{ \omega, \epsilon_{n,\ell}^{(m)} f_{n,\ell}(\Theta) + \epsilon_{n,\ell}^{(a)} \right\}$$

Θ

parameters to infer

$f_{n,\ell}$

black-box, spans **multiple decades** (element (n, ℓ))

$$\epsilon_{n,\ell}^{(a)} \sim \mathcal{N}(0, \sigma_a^2)$$

e.g., instruments noise

$$\epsilon_{n,\ell}^{(m)} \sim \log \mathcal{N}(0, \sigma_m^2)$$

e.g., calibration error

$$\omega > 0$$

instrument detectability limit

How to deal with

black-box forward map f ?

mixture of **additive** and **multiplicative** noises?

How to deal with **black-box** forward map f ?

→ Model reduction

How to deal with **mixture** of **additive** and **multiplicative** noises?

→ likelihood approximation with controlled error

Deriving a new Likelihood (uncensored)

$$Y = \epsilon^{(m)} f(\Theta) + \epsilon^{(a)}$$

	Additive approx	Multiplicative approx
Approximated model	$Y \simeq f(\Theta) + e^{(a)}$	$Y \simeq e^{(m)} f(\Theta)$
Moment matching	$e^{(a)} \sim \mathcal{N}(\mathbf{m}_a, \mathbf{s}_a^2)$	$e^{(m)} \sim \log \mathcal{N}(\mathbf{m}_m, \mathbf{s}_m^2)$
likelihood approx (1 elt)	$\pi^{(a)}(y_{n,\ell} \Theta)$	$\pi^{(m)}(y_{n,\ell} \Theta)$

Deriving a new Likelihood (uncensored)

$$\tilde{\pi}(y_{n,\ell}|\Theta) \propto \pi^{(a)}(y_{n,\ell}|\Theta)^{1-\lambda_{n,\ell}} \pi^{(m)}(y_{n,\ell}|\Theta)^{\lambda_{n,\ell}}$$

with $\lambda_{n,\ell} = \lambda(f_{n,\ell}(\Theta))$: | controls the mixing of the two approx
sigmoid parametrized by \mathbf{a}_ℓ (location and speed)

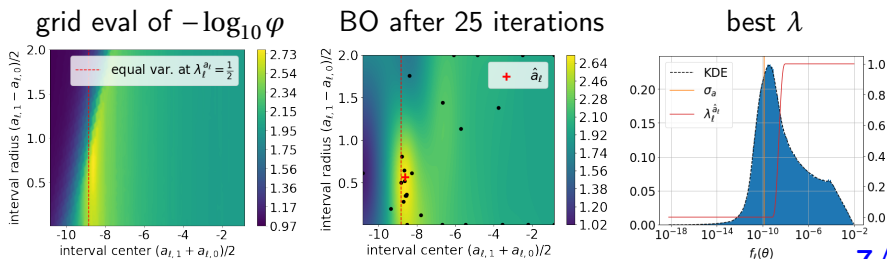
Deriving a new Likelihood (uncensored)

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with $\lambda_{n,\ell} = \lambda(f_{n,\ell}(\Theta))$: controls the mixing of the two approx sigmoid parametrized by \mathbf{a}_ℓ (location and speed)

To evaluate approx: Kolmogorov-Smirnov-based metric $\varphi(\mathbf{a}_\ell)$.

To get best approx: Minimize φ with Bayesian Optimization (BO).



A priori information

a priori information on $\Theta \in \mathbb{R}^{N \times D}$ combines 2 priors:

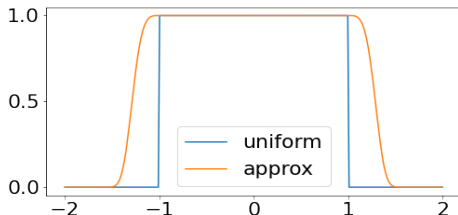
- **spatial regularization**, e.g.,

- smoothed Total Variation
- L_2 -norm of image gradient
- L_2 -norm of image Laplacian
- L_2 -norm of image wavelet decomposition

- **Validity domain** for each physical parameter $\theta_{n,d}$

⇒ BUT non-smooth

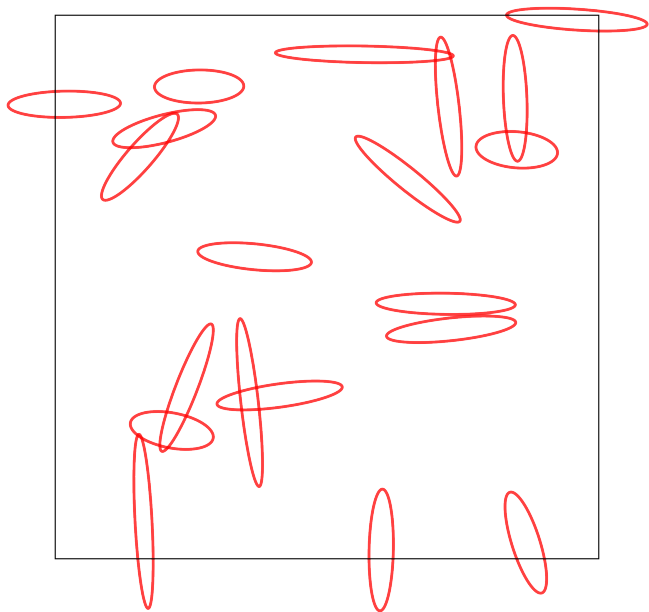
⇒ use smooth penalty function when $\theta_{n,d}$ is out of validity domain:



Proposed sampler: **two kernels**:

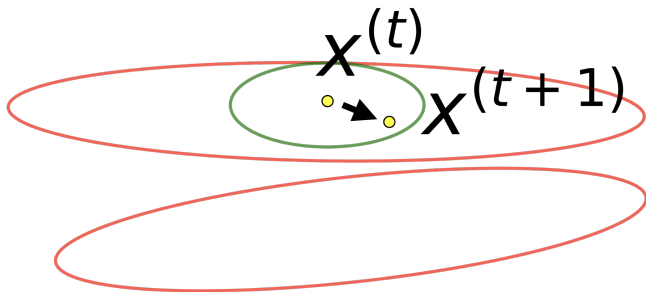
- Forward model covers **multiple decades**
⇒ non-grad.-Lipschitz log-posterior
Preconditioned-MALA kernel with RMSProp
Role: Efficient local exploration
Limitation: restricted to smooth log-posteriors
- **Non-log-concave** posterior
⇒ potential multimodality
Multiple-Try Metropolis (MTM) kernel
Role: Jumps between modes

Sampler: illustration on Gaussian mixture model

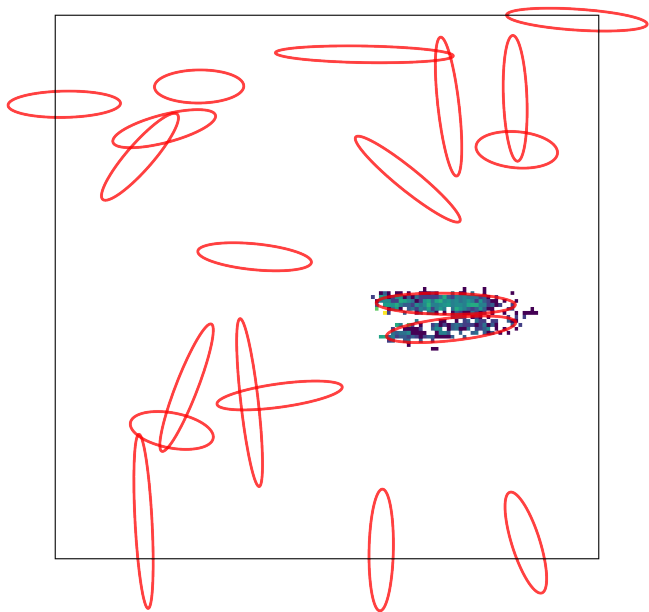


Sampler: illustration on Gaussian mixture model

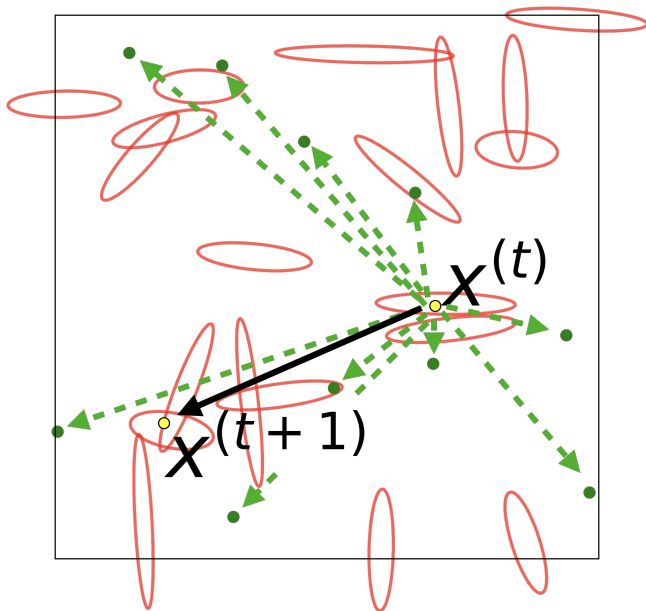
 proposal distribution covariance



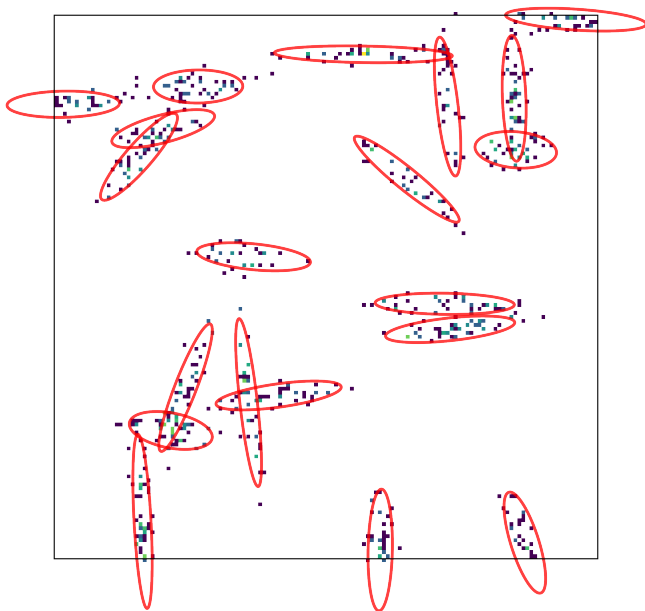
Sampler: illustration on Gaussian mixture model



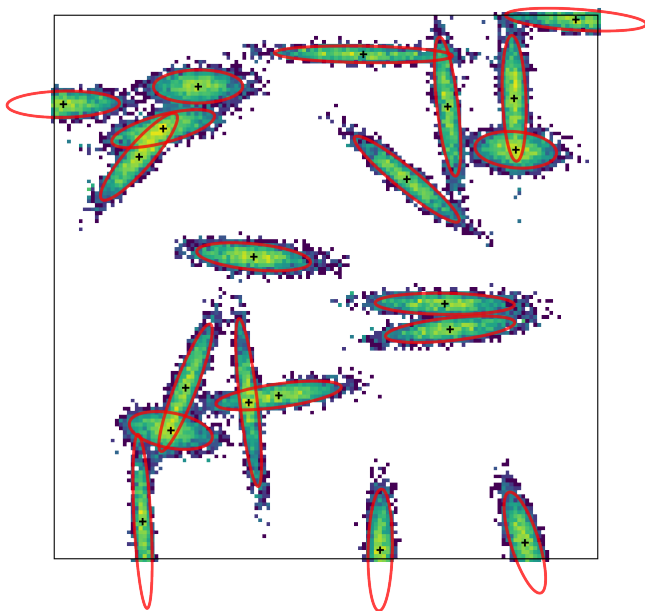
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Sampler: illustration on Gaussian mixture model

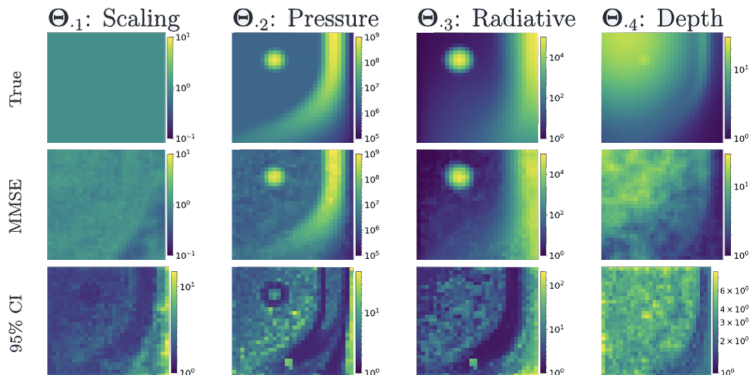
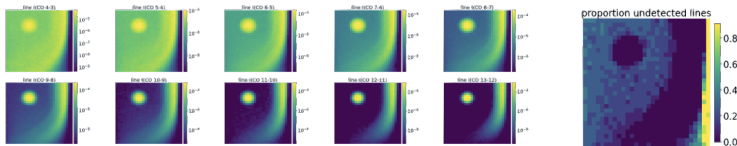


Sampler: illustration on Gaussian mixture model



Application to an astrophysics synthetic dataset

Synthetic observations $\mathbf{Y} \in \mathbb{R}^{900 \times 10}$: integrated intensities of excited lines of CO



Conclusion

Log-likelihood

Untractable

Non-concave

Non-grad-Lipschitz

Proposition

Approx.

Model reduction

MTM kernel

P-MALA kernel

Mixture of noises

Forward model:

→ Black-box expensive

→ Non-linear

→ Spans multiple decades

Applications ✓ Astrophysics synthetic yet realistic dataset

→ OrionB data and JWST obs. (perspective)

